

## 6.6. Dirac Electrons in Magnetic Field

In the presence of a static magnetic field, the Dirac Hamiltonian becomes (see §6.1)

$$H = c \begin{pmatrix} mc & \boldsymbol{\sigma} \cdot \mathbf{P} \\ \boldsymbol{\sigma} \cdot \mathbf{P} & -mc \end{pmatrix}$$

where

$$\mathbf{P} = \mathbf{p} - \frac{q}{c} \mathbf{A}$$

For a uniform  $\mathbf{B} = B \hat{\mathbf{z}}$ , the simplest solutions to

$$\nabla \times \mathbf{A} = \begin{vmatrix} \hat{x} & \hat{y} & \hat{z} \\ \partial_x & \partial_y & \partial_z \\ A_x & A_y & A_z \end{vmatrix} = (0, 0, B)$$

are those with  $A_z = 0$ , e.g.,

$$\mathbf{A} = (0, xB, 0) \quad \text{or} \quad \mathbf{A} = \frac{1}{2}(-yB, xB, 0)$$

If the particles are also confined to the xy-plane, we have

$$\boldsymbol{\sigma} \cdot \mathbf{P} = \begin{pmatrix} 0 & P_- \\ P_+ & 0 \end{pmatrix} \equiv Q$$

where

$$P_{\pm} = P_x \pm iP_y$$

$$\begin{aligned} [P_x, P_y] &= \left[ p_x - \frac{q}{c} A_x, p_y - \frac{q}{c} A_y \right] \\ &= \left[ p_x, -\frac{q}{c} A_y \right] + \left[ -\frac{q}{c} A_x, p_y \right] \\ &= i\hbar \frac{q}{c} (\partial_x A_y - \partial_y A_x) \\ &= i\hbar \frac{q}{c} B \end{aligned}$$

$$\begin{aligned} \rightarrow [P_+, P_-] &= [P_x, -iP_y] + [iP_y, P_x] \\ &= -2i[P_x, P_y] \\ &= 2\hbar \frac{q}{c} B \end{aligned}$$

Let

$$a = \sqrt{\frac{c}{2\hbar qB}} P_+ \quad a^+ = \sqrt{\frac{c}{2\hbar qB}} P_-$$

$$\rightarrow [a, a^+] = 1$$

$$\& Q = \sqrt{\frac{2\hbar qB}{c}} \begin{pmatrix} 0 & a^+ \\ a & 0 \end{pmatrix} = \begin{pmatrix} 0 & A^+ \\ A & 0 \end{pmatrix}$$

$$\text{where} \quad A = \sqrt{\frac{2\hbar qB}{c}} a$$

$$\begin{aligned} \therefore Q^2 &= \begin{pmatrix} 0 & A^+ \\ A & 0 \end{pmatrix} \begin{pmatrix} 0 & A^+ \\ A & 0 \end{pmatrix} = \begin{pmatrix} A^+A & 0 \\ 0 & AA^+ \end{pmatrix} \\ &= \frac{2\hbar qB}{c} \begin{pmatrix} a^+a & 0 \\ 0 & aa^+ \end{pmatrix} \end{aligned}$$

The associated Fock states

$$|N\rangle = \frac{1}{\sqrt{N!}} (a^+)^N |0\rangle$$

are called the  $N^{\text{th}}$  Landau levels:

Also

$$H = c \begin{pmatrix} mc & Q \\ Q & -mc \end{pmatrix}$$

$$\begin{aligned} \rightarrow H^2 &= c^2 \begin{pmatrix} mc & Q \\ Q & -mc \end{pmatrix} \begin{pmatrix} mc & Q \\ Q & -mc \end{pmatrix} \\ &= c^2 \begin{pmatrix} (mc)^2 + Q^2 & 0 \\ 0 & (mc)^2 + Q^2 \end{pmatrix} \end{aligned}$$

$$\therefore H = c \begin{pmatrix} \sqrt{(mc)^2 + Q^2} & 0 \\ 0 & -\sqrt{(mc)^2 + Q^2} \end{pmatrix}$$

Note that both  $(\pm)$  roots of square root  $(mc)^2 + Q^2$  are needed to ensure all eigenvalues of  $H$  are covered.

In the non-relativistic limit,

$$c\sqrt{(mc)^2 + Q^2} \approx mc^2 + \frac{Q^2}{2m}$$

$$\rightarrow H \approx \frac{Q^2}{2m} \begin{pmatrix} 1 & 0 \\ 0 & -1 \end{pmatrix}$$

while

$$\frac{Q^2}{2m} = \hbar \frac{qB}{mc} \begin{pmatrix} a^+a & 0 \\ 0 & aa^+ \end{pmatrix} = \hbar\omega_c \begin{pmatrix} a^+a & 0 \\ 0 & aa^+ \end{pmatrix}$$

where

$$\omega_c = \frac{qB}{mc}$$

is the cyclotron frequency.

Setting

$$k_Q = \sqrt{\frac{2m\omega_c}{\hbar}} = \sqrt{\frac{2qB}{\hbar c}}$$

we have

$$Q^2 = (\hbar k_Q)^2 \begin{pmatrix} a^+a & 0 \\ 0 & aa^+ \end{pmatrix}$$

Thus

$$\langle N | Q^2 | N \rangle = (\hbar k_Q)^2 \begin{pmatrix} N & 0 \\ 0 & N+1 \end{pmatrix}$$

$$\langle N | \sqrt{(mc)^2 + Q^2} | N \rangle = \frac{1}{c} \begin{pmatrix} \mathcal{E}(N) & 0 \\ 0 & \mathcal{E}(N+1) \end{pmatrix}$$

where

$$\mathcal{E}(N) = c \sqrt{(mc)^2 + N(\hbar k_Q)^2}$$

$$\therefore \langle N | H | N \rangle = \begin{pmatrix} \mathcal{E}(N) & 0 & 0 & 0 \\ 0 & \mathcal{E}(N+1) & 0 & 0 \\ 0 & 0 & -\mathcal{E}(N) & 0 \\ 0 & 0 & 0 & -\mathcal{E}(N+1) \end{pmatrix}$$

The eigen-spinors are simply

$$u_{\uparrow}^{\mathcal{E}(N)} = \begin{pmatrix} 1 \\ 0 \\ 0 \\ 0 \end{pmatrix} \quad u_{\downarrow}^{\mathcal{E}(N+1)} = \begin{pmatrix} 0 \\ 1 \\ 0 \\ 0 \end{pmatrix} \quad v_{\uparrow}^{-\mathcal{E}(N)} = \begin{pmatrix} 0 \\ 0 \\ 1 \\ 0 \end{pmatrix} \quad v_{\downarrow}^{-\mathcal{E}(N+1)} = \begin{pmatrix} 0 \\ 0 \\ 0 \\ 1 \end{pmatrix}$$

Thus, associated with a given Landau level  $|N\rangle$ , we have 2 energy eigenstates of positive energies:  $\mathcal{E}(N)$  for spin  $\uparrow$  particles &  $\mathcal{E}(N+1)$  for spin  $\downarrow$  particles.

Conversely, a given energy level  $\mathcal{E}(N)$ , except for  $N=0$ , can belong to either a spin  $\uparrow$  particle in Landau level  $|N\rangle$  or a spin  $\downarrow$  particle in  $|N-1\rangle$ .

For the vacuum  $|0\rangle$ , we have a spin  $\uparrow$  particle with energy  $\mathcal{E}(0) = mc^2$ , & a spin  $\downarrow$  particle with energy  $\mathcal{E}(1) = c \sqrt{(mc)^2 + (\hbar k_Q)^2}$ . (see Ezawa, Fig.6.2)

Comparing the Hamiltonian

$$H_Q = c^2 Q^2 = c^2 (\boldsymbol{\sigma} \cdot \mathbf{P})^2$$

with the Pauli Hamiltonian (see 6.4.a.\_NonrelativisticLimit.pdf)

$$H_P = \frac{1}{2m} (\boldsymbol{\sigma} \cdot \mathbf{P})^2 = \frac{1}{2m} \left( \mathbf{p} - \frac{q}{c} \mathbf{A} \right)^2 - \frac{q\hbar}{2mc} \boldsymbol{\sigma} \cdot \mathbf{B}$$

$$= \frac{1}{2m} \left( \mathbf{p} - \frac{q}{c} \mathbf{A} \right)^2 - \frac{1}{2} \hbar \omega_c \boldsymbol{\sigma} \cdot \hat{\mathbf{B}} \quad \left( \omega_c = \frac{qB}{mc} \right)$$

we see that  $H_Q$  is just a Pauli Hamiltonian with an effective mass  $m_Q = \frac{1}{2c^2}$ .

$$\rightarrow H_Q = c^2 \left( \mathbf{p} - \frac{q}{c} \mathbf{A} \right)^2 - q\hbar c \boldsymbol{\sigma} \cdot \mathbf{B}$$

Note that the Zeeman energies of  $H_P$  are  $\pm \frac{1}{2} \hbar \omega_c$ ,

while the cyclotron energy is  $\hbar \omega_c$ .

Thus, the spectrum  $\pm \mathcal{E}(N)$  of the relativistic Dirac Hamiltonian  $H$  are constructed from the spectrum

$$E(N) = N(\hbar c k_Q)^2 = 2q\hbar c B N$$

of the non-relativistic Hamiltonian  $H_Q$ .

For  $H_Q$ , the Zeeman effect associated with the  $\boldsymbol{\sigma} \cdot \mathbf{B}$  term is called intrinsic since  $\boldsymbol{\sigma}$  is a built-in property of the Dirac Hamiltonian.

The intrinsic Zeeman energies are  $E_Z = \pm q\hbar c B$ .

The cyclotron energy in  $H_Q$  is

$$\hbar \omega_c^Q = \hbar \frac{qB}{m_Q c} = 2q\hbar c B$$

Caution: These “energies” have dimensions of  $(\text{energy})^2$  since we’ve set  $m_Q c^2 = \frac{1}{2}$  dimensionless.

On the other hand, associated with  $k_Q = \sqrt{\frac{2qB}{\hbar c}}$  is a frequency

$$\omega_Q = k_Q c$$

so that  $\mathcal{E}(N) = \sqrt{(m c^2)^2 + N(\hbar \omega_Q)^2}$ .

Note that  $\hbar \omega_Q = \hbar k_Q c = \sqrt{2q\hbar c B} = \sqrt{\hbar \omega_c^Q}$ .